

Joint Institute for Nuclear Research



Welcome to Dubna

Baksan School on AstroParticle Physics
April 17, 2019

Dmitry V.Naumov

Dubna



Moscow sea

Volga River

Dubna River

Moscow Canal

Sister River

A magic island

Dubna. JINR. DLNP Campus



Dubna. JINR. VBLHE Campus



JINR

- New elements 102, {103, 104, 105(Db), 107}, 114, 115, 116, 117, 118
- Hypothesis of neutrino oscillations (1957). {NP: 2015}
- Discovery of new particle: anti-sigma-minus hyperon
- And many other discoveries



JINR

- Employed ~ 5000: 1200 - scientists, 2000 - engineers
- 7 labs. Each as a big institute
- 18 countries member-states and 6 associated
- 1500 science papers/year
- Collaboration with 800 scientific centers in 64 countries
- 1, 5 billions USD. Budget for 2017-2023
- Laboratory of high energy physics
- Laboratory of nuclear problems
- Laboratory of nuclear reactions
- Laboratory of neutron physics
- Laboratory of information technologies
- Laboratory of theoretical physics
- Laboratory of radiation biology

JINR

Short History

History

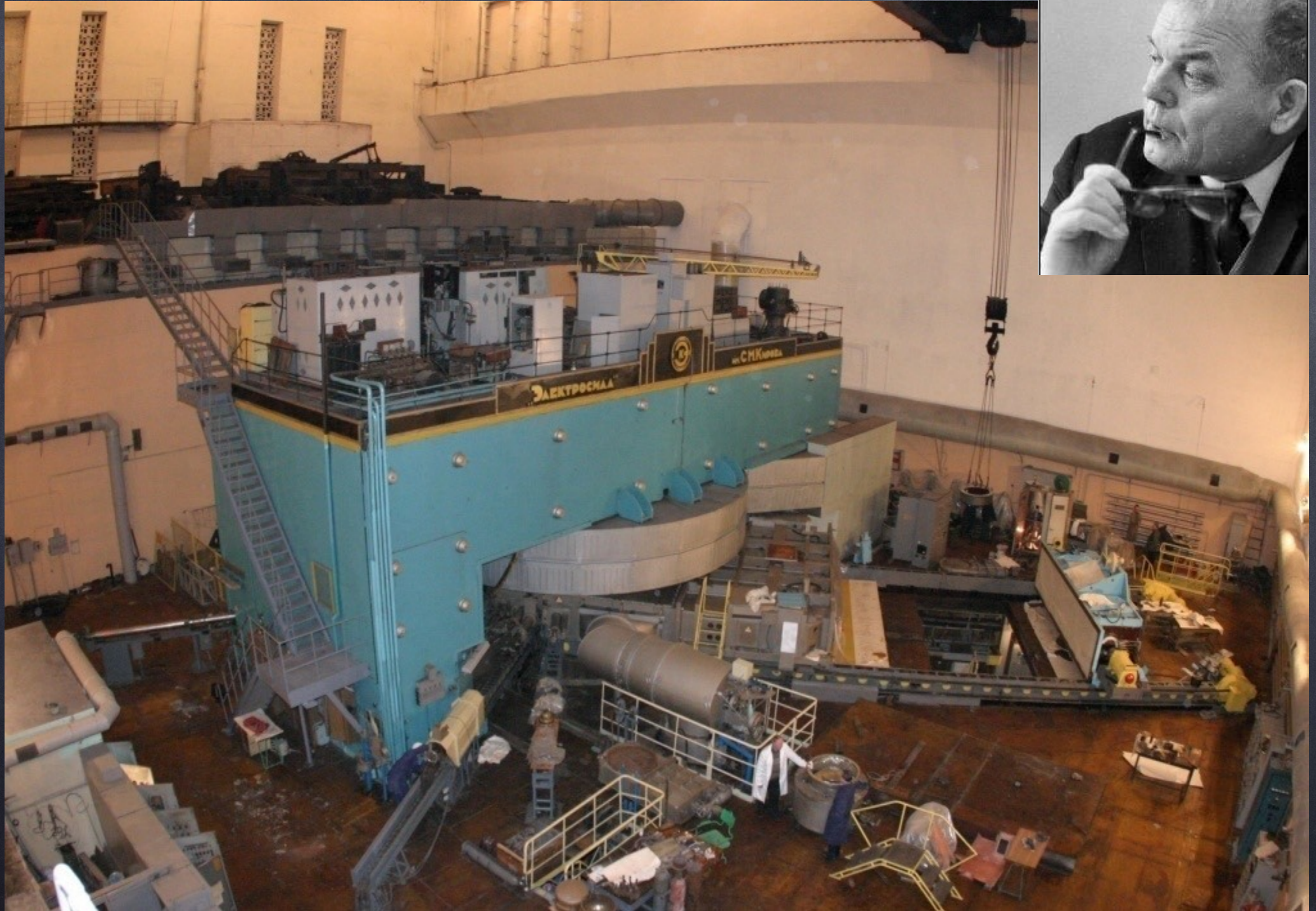
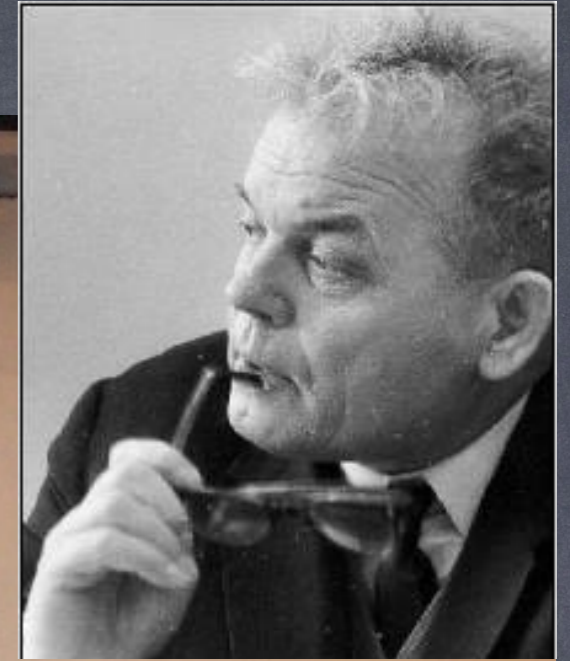
- May, 7 1946. First discussion of «construction of a power cyclotron» at special committee of the government
- 18 August 1946. Soviet government approved the proposal of Academician Igor Kurchatov to construct in USSR „the installation M” for fundamental studies in nuclear physics.
- 14 December 1949. The 480 MeV proton synchrocyclotron started operation at the Hydrotechnical Laboratory in Dubna, the most powerful accelerator in the world at that time.
- Electrophysical lab lead by Veksler worked on proton accelerator with 10 GeV energy
- 26 March 1956. On the basis of two labs JINR founded.

Synchrocyclotron of DLNP JINR



Synchrocyclotron. 680 MeV (1953)

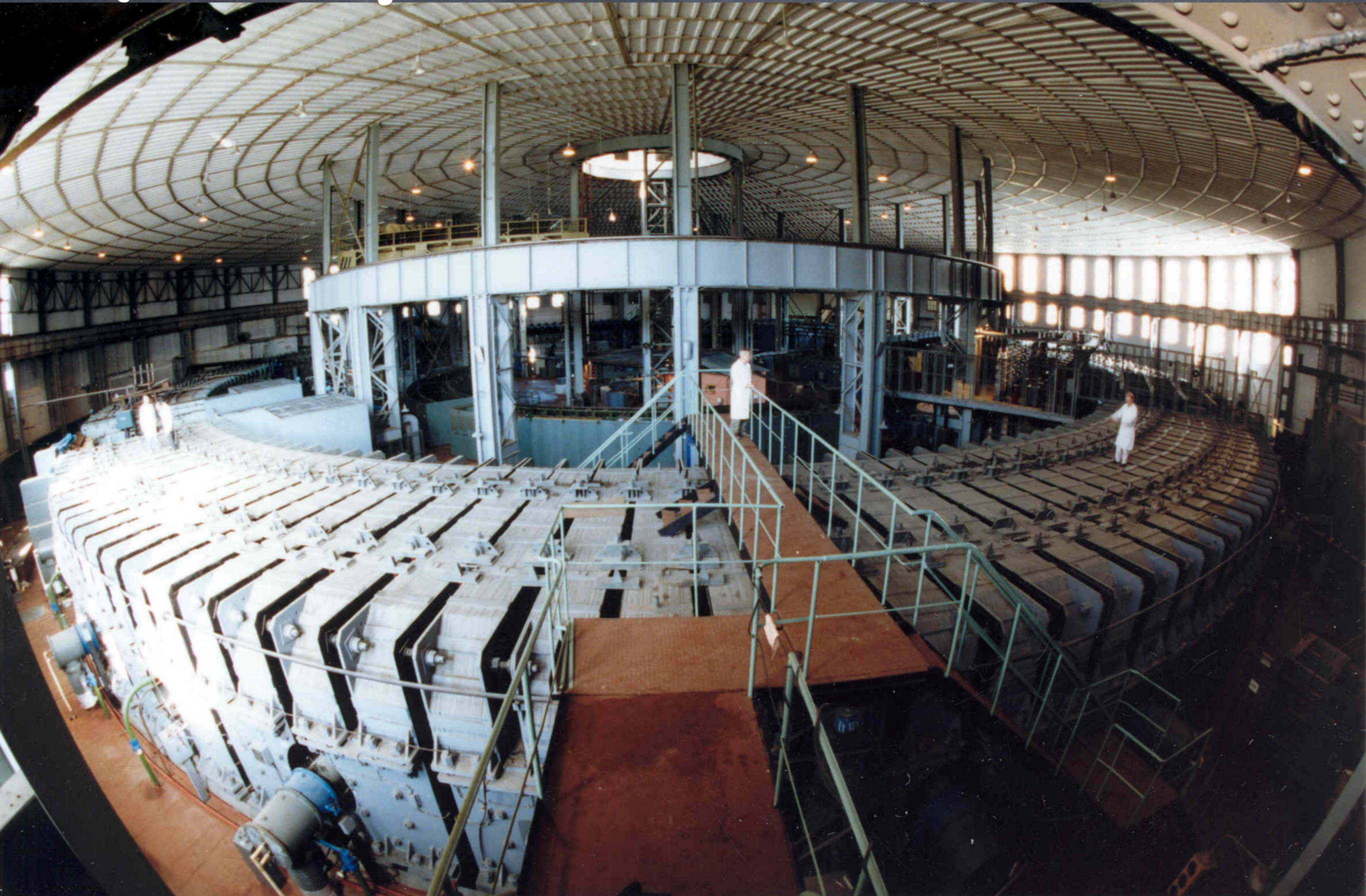
M.G. Mecheriakov



Synchrotron. VBLHEP JINR



Synchrotron. VBLHEP JINR



Two facts

- Half of discoveries in nuclear physics belongs to JINR (37)
- 15 belongs to DLNP JINR

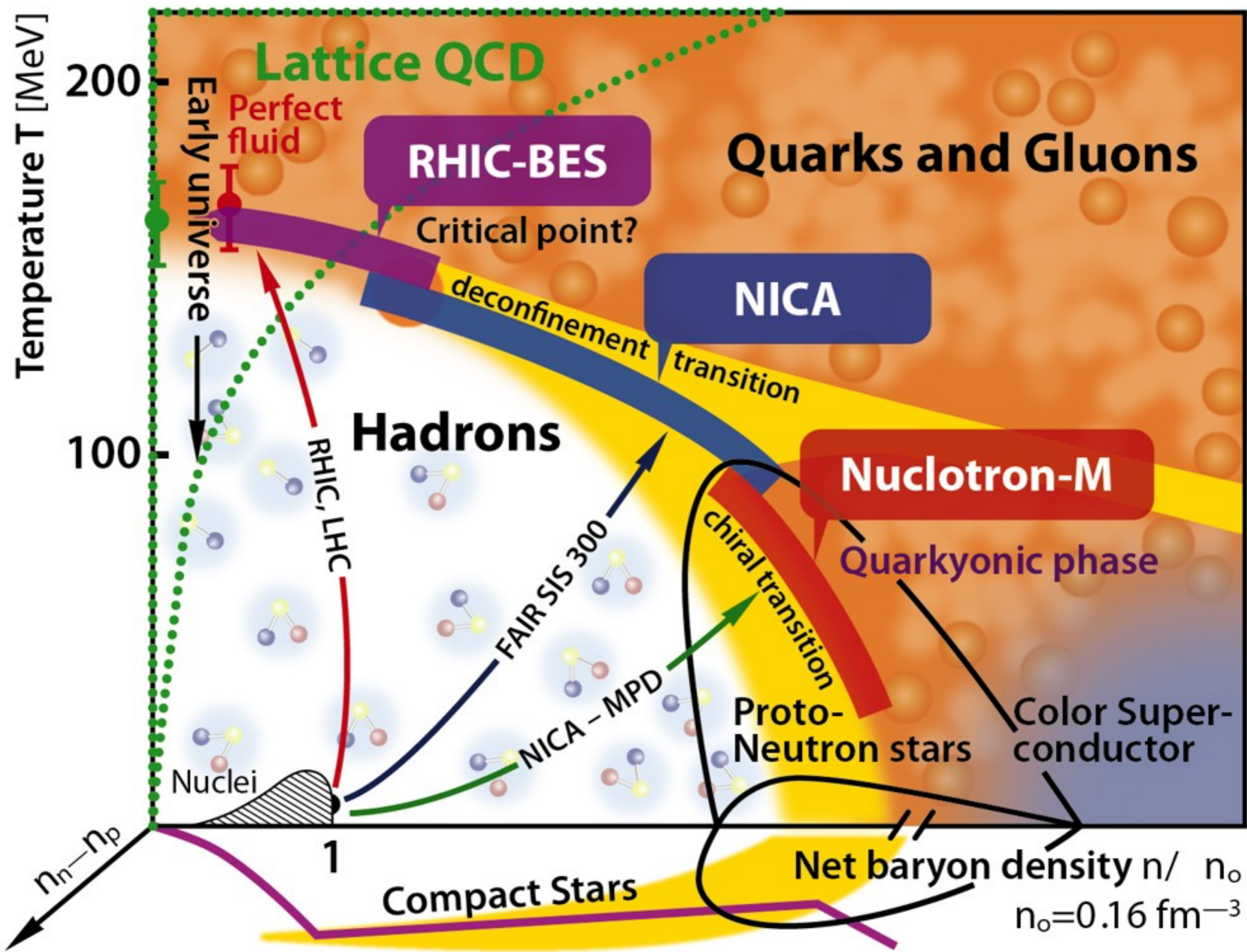
VBLHEP

Veksler and Baldin Laboratory of High Energy Physics

http://lhe.jinr.ru/index_rus.html

NICA - Nuclotron-based Ion Collider fAcility
Flagship project of VBLHEP and JINR
Mega-Science project of Russia





Observables

Sensitive to phase transitions

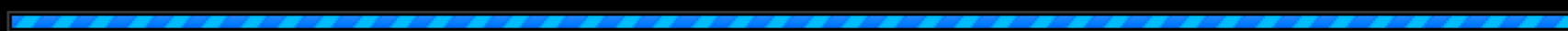
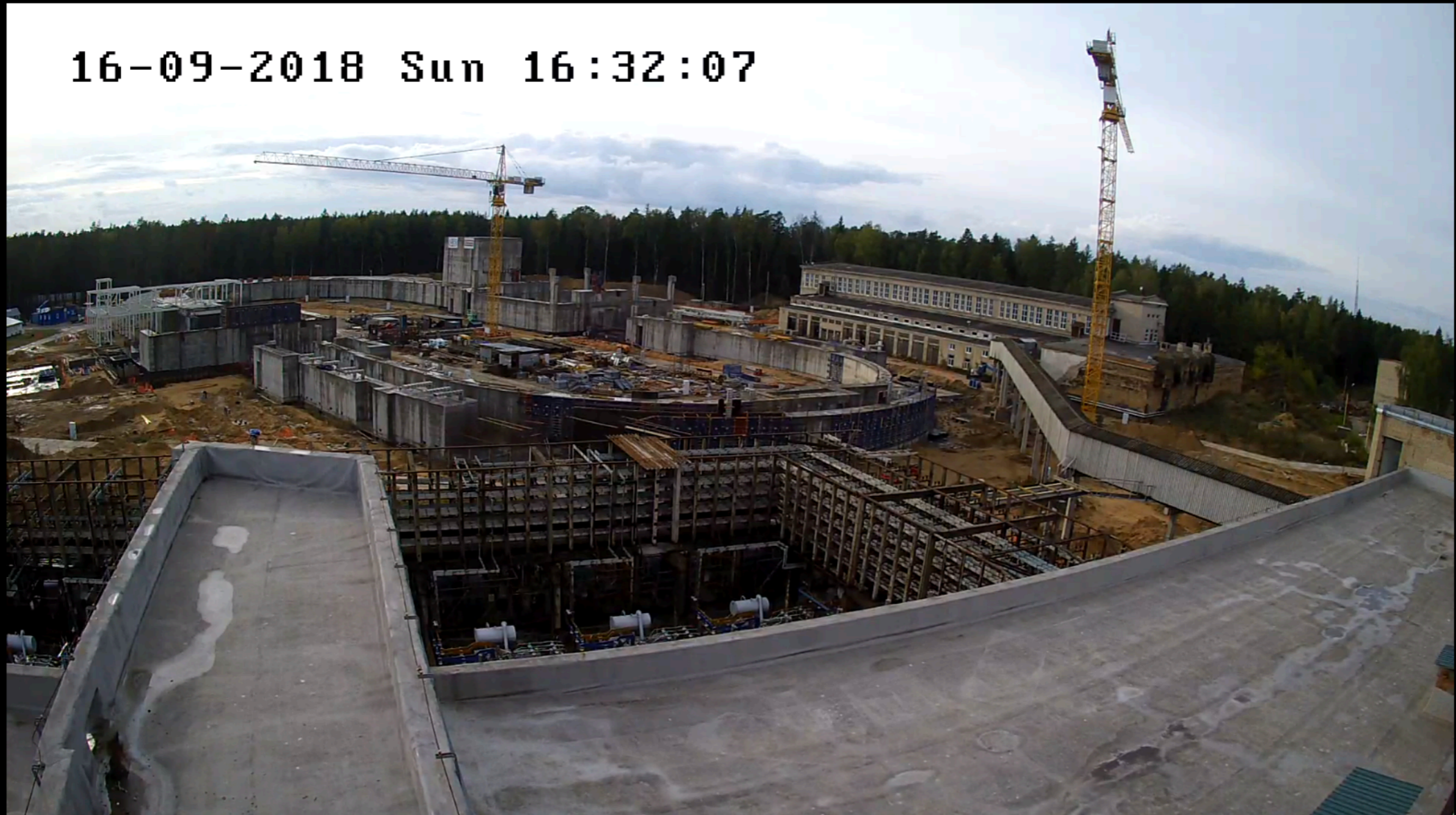
- Particle's yields and spectra
- Event-by-event fluctuations of multiplicity and p_T
- Restoration of chiral symmetry \rightarrow modification of hadronic spectral functions
- Strangeness production enhancement

Строительство коллайдера NICA

<http://nucloweb.jinr.ru/nucloserv/205corp.htm>



16-09-2018 Sun 16:32:07

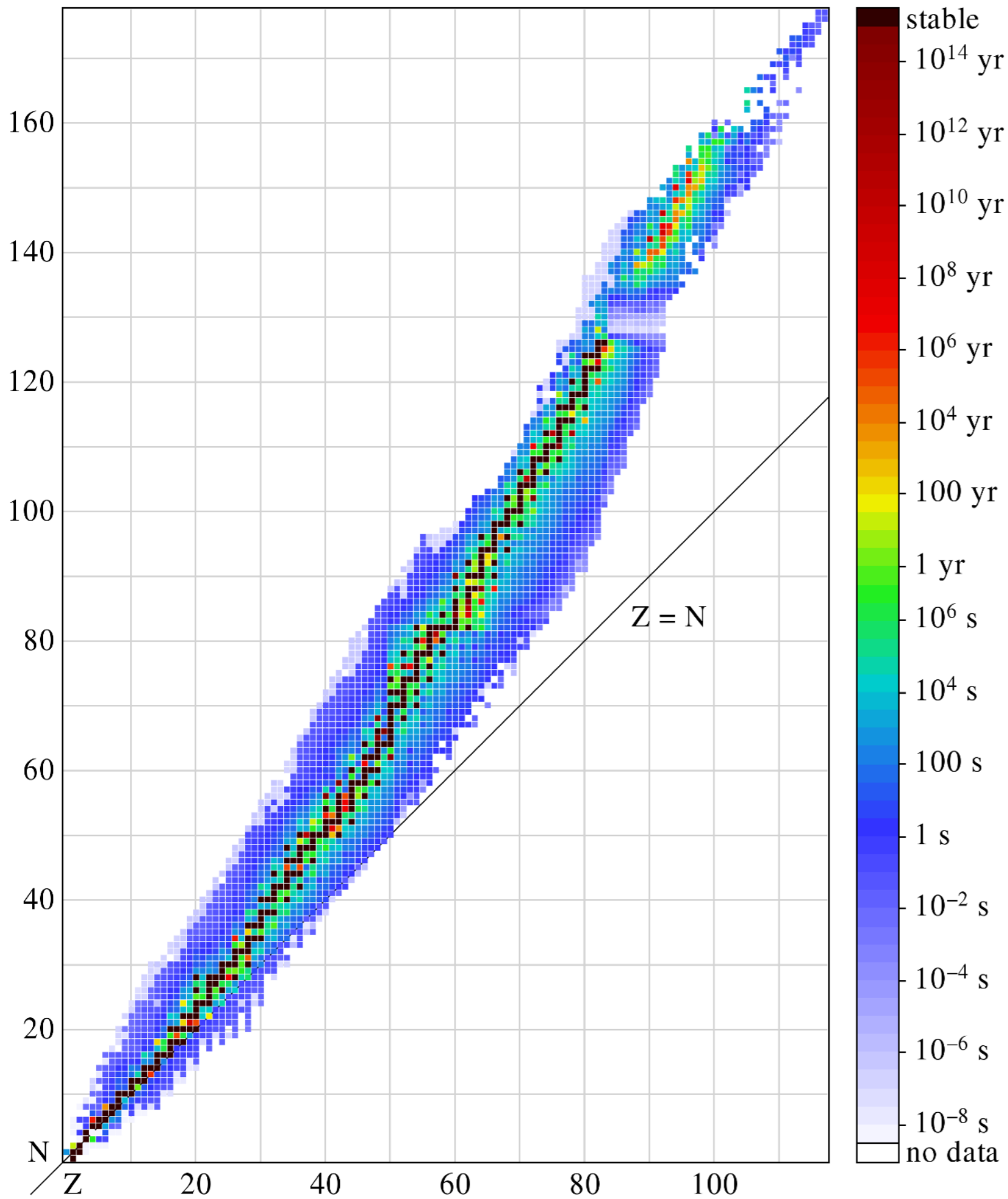


Live



FLNR

Flerov Laboratory of Nuclear Reactions



• Nuclei become less stable with increasing Z and N

• Shell model \rightarrow new stability island with $Z=114$ and $N=184$

Stable 'Mountains'
Lead - Uranium

Spontaneous Fission

Island of Stability
- superheavy spherical nuclei

SEA OF INSTABILITY

100

130

120

110

100

90

80

70

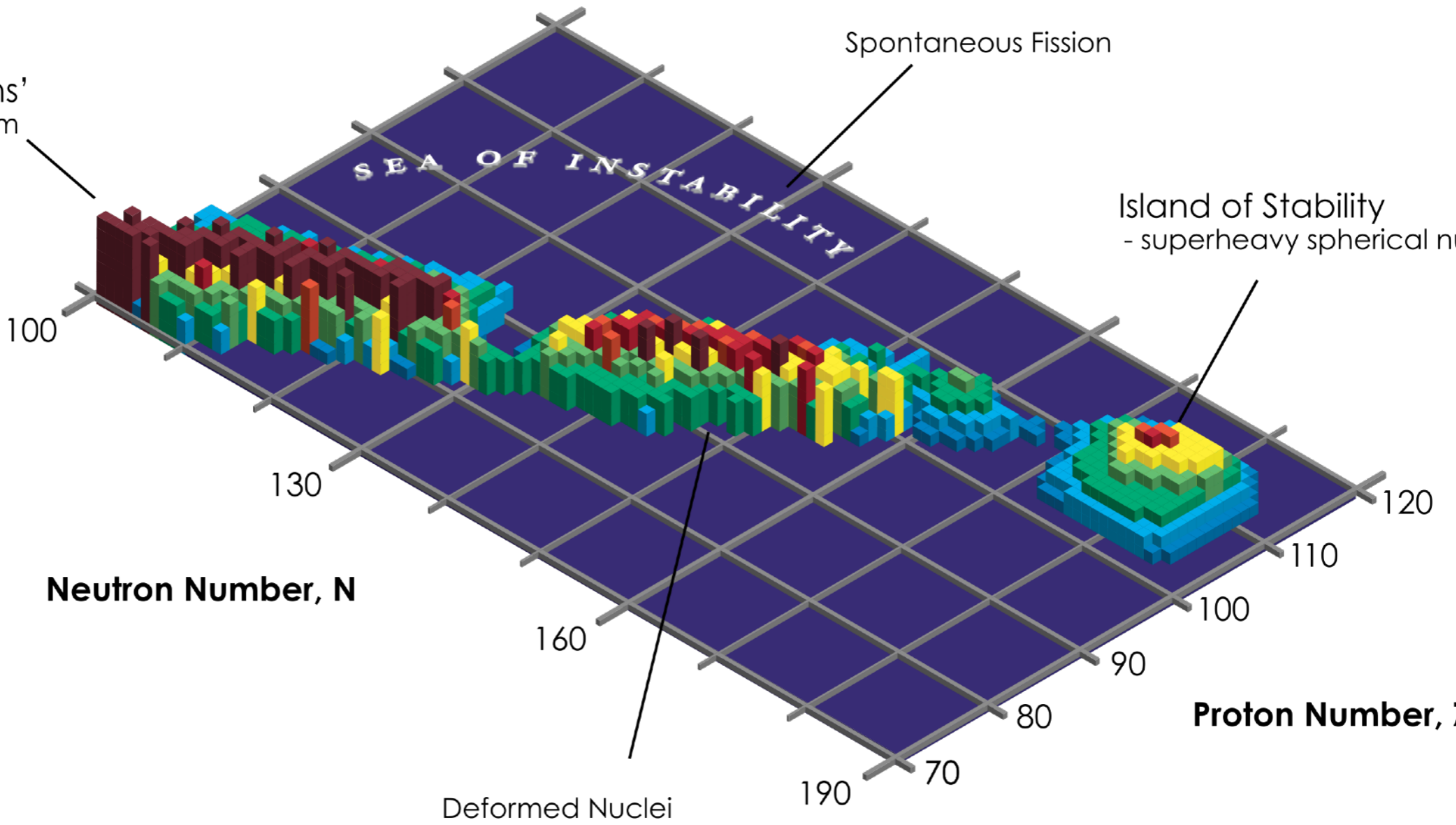
190

Neutron Number, N

Proton Number, Z

Deformed Nuclei

Increasing Stability ↑



Elements synthesized in Dubna

Нобелий 102 _{5f¹⁴} No [259] Nobelium	Лоуренсий 103 _{5f¹⁴6d¹} Lr [266] Lawrencium	Резерфордий 104 _{6d²} Rf [261] Rutherfordium	Дубний 105 _{6d³} Db [268] Dubnium	Хассий 108 _{6d⁸} Hs [269] Hassium
Флеровий 114 Fl [289] Flerovium	Московский 115 Mc [290] Moscovium	Ливерморий 116 Lv [293] Livermorium	Теннессин 117 Ts [294] Tennessine	Оганесон 118 Og [294] Oganesson

Z=114

- The exact location of stability is not yet found
- The trend of increasing stability with $N \rightarrow 184$ is observed

SHE: Super-Heavy Elements factory

DC-280
Фабрика
сверхтяжелых
элементов



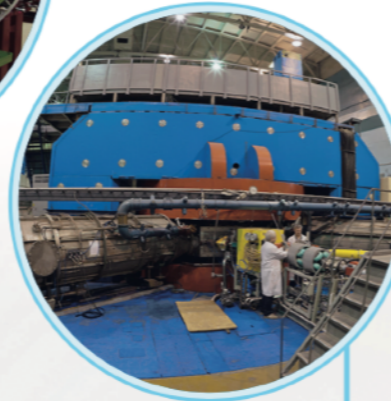
U-400
Тяжелые и
сверхтяжелые
ядра



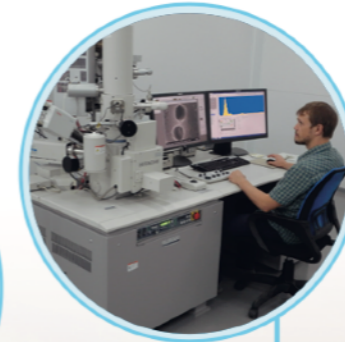
ACCULINNA-2
Фрагмент-сепаратор



U-400M
Лёгкие
экзотические
ядра

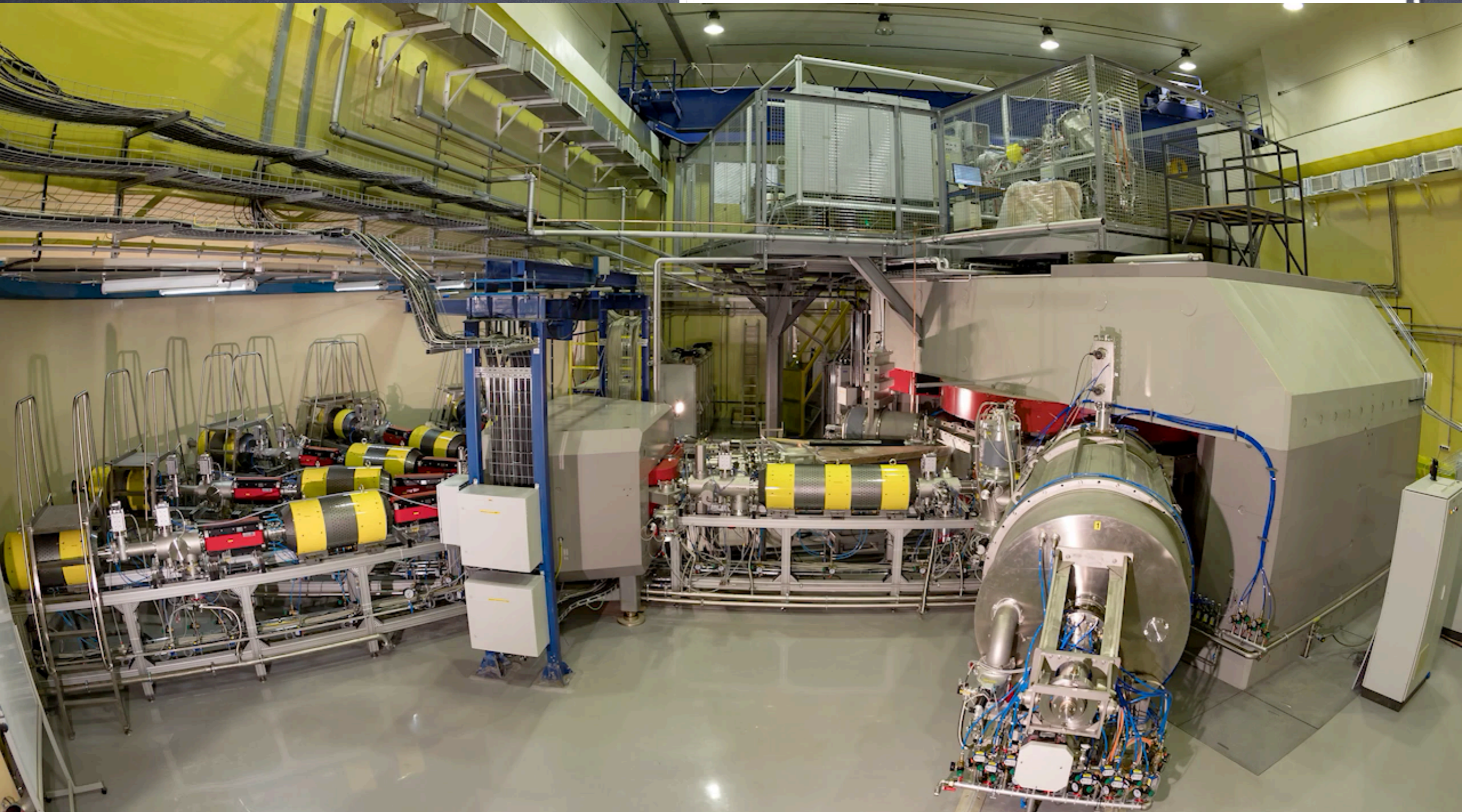


**Нано-
лаборатория**



Кому **Ruthenium. Moscovium.**
Dubnium. Oganesson

Куда _____



FLNP

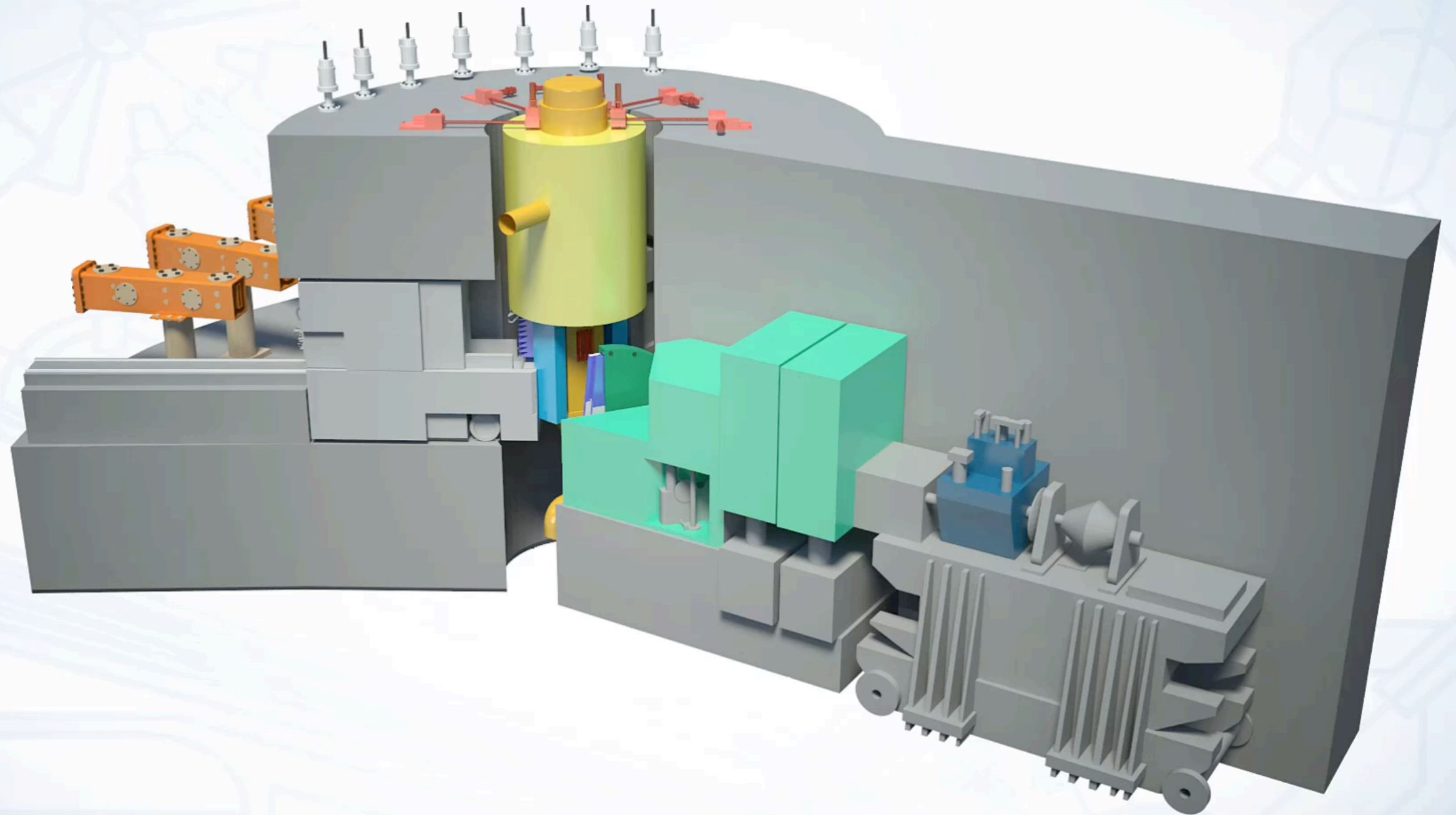
Frank Laboratory of Neutron Physics

<http://flnph.jinr.ru/en/>

Impulse fast reactor (IBR)

- End of 1955. Idea by D.I. Blokhintsev during a seminar in Obninsk
- Mid of 1956. D.I. Blokhintsev was suggested to head JINR. His condition was — build IBR
- November 1960. I.M. Frank reported first results from newly constructed IBR.
- ...
- 2010. Upgraded IBR-2

IBR-2



R&D for new most powerful neutron source

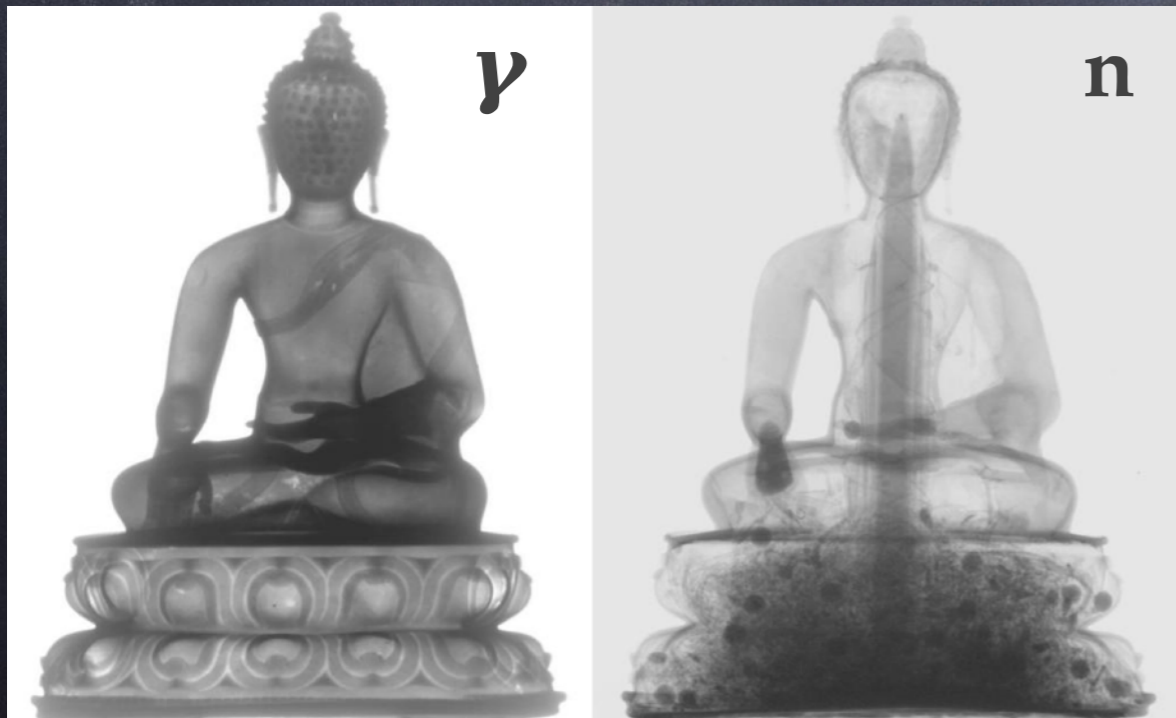
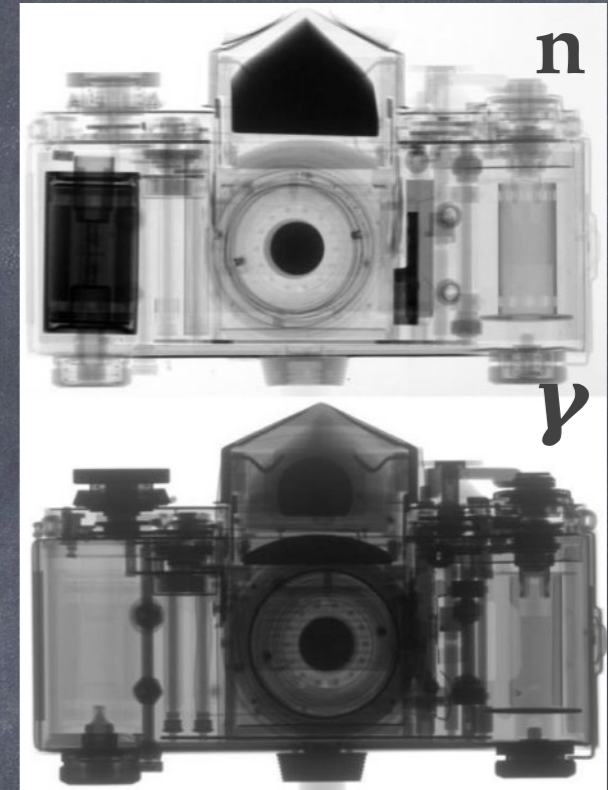
Wide research program

- Study neutron
- Study structure and dynamics of condensed matter
 - Crystals, nanosystems
 - Complex liquids, polymers, rock
 - Molecular biology and pharmacology
- Applied research

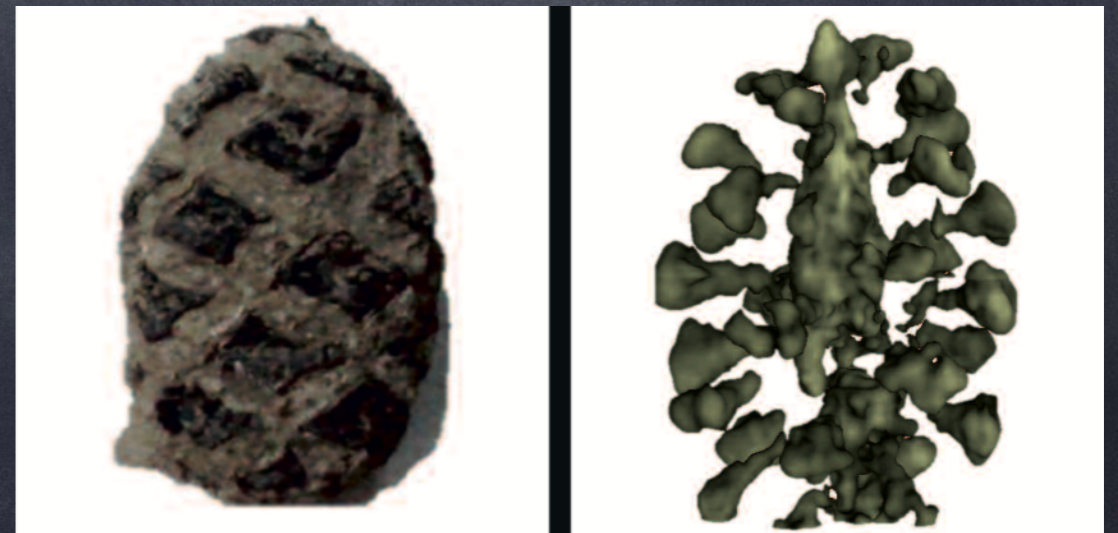
Neutron radiography



- Organic structure is better seen with help of neutrons
- Gamma is sensitive to density only

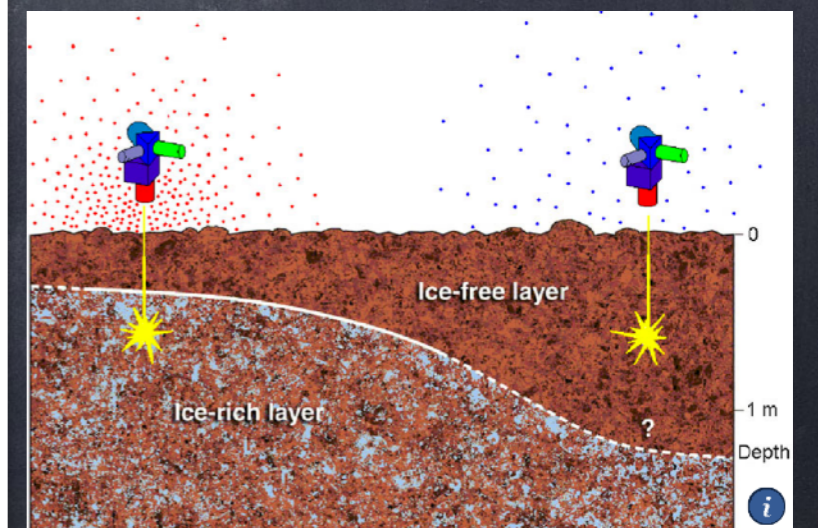


Effective for paleontology



Search for water on Mars

- Mars Odyssey 2001 is on Mars orbit from 2002.
- Neutron detector HEND (with help of JINR) is onboard.
- Source and detector of neutrons DAN (JINR) is onboard of Mars rover «Curiosity Rover»



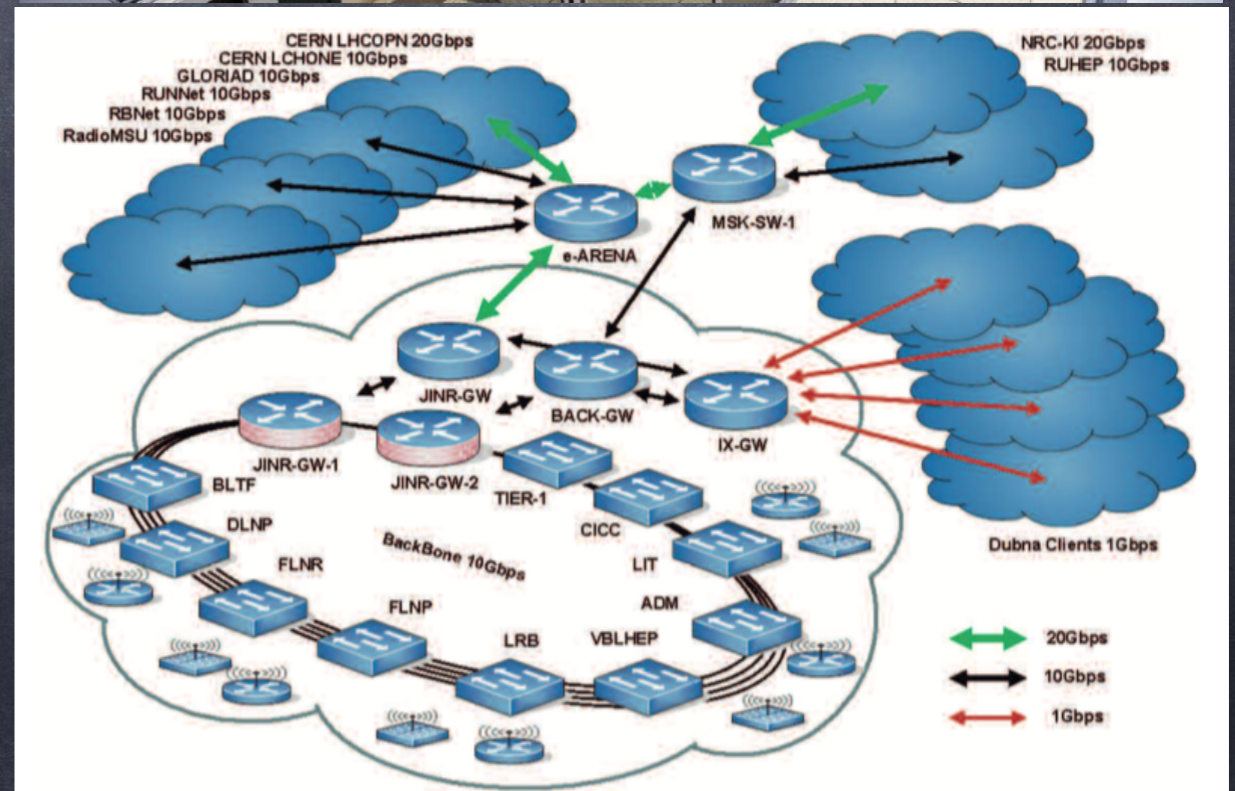
LIT

Laboratory of information technologies

<http://lit.jinr.ru>

Computing cluster

- Included in computing network GRID
- Used by leading centers: CERN, IHEP, Fermilab...
- More than 6000 CPU units
- Heterogeneous cluster (GPU)
- 4.2 petabytes of disk space
- 5 petabytes on robotic tapes
- Support 24/7



BLTP

Bogoliubov Laboratory of Theoretical Physics

http://theor.jinr.ru/lab_rus.shtml

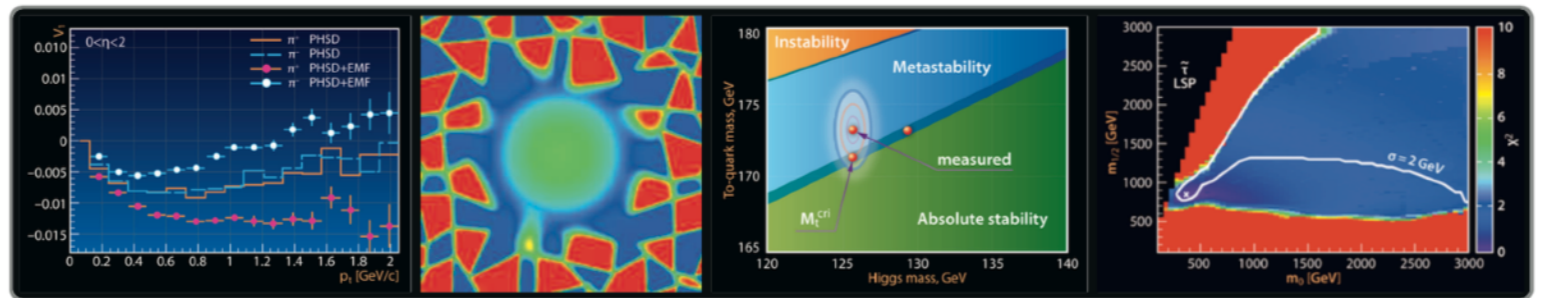
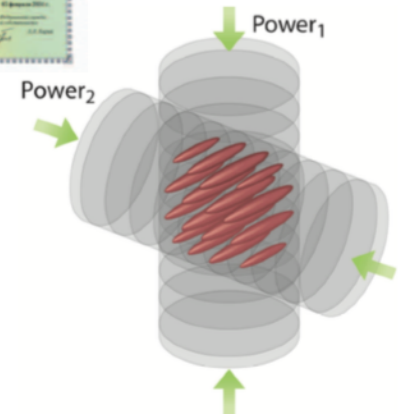
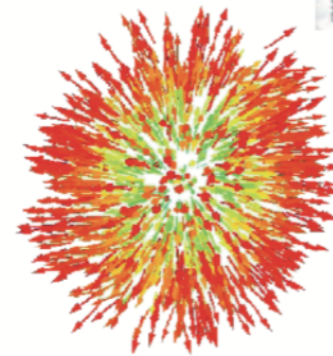
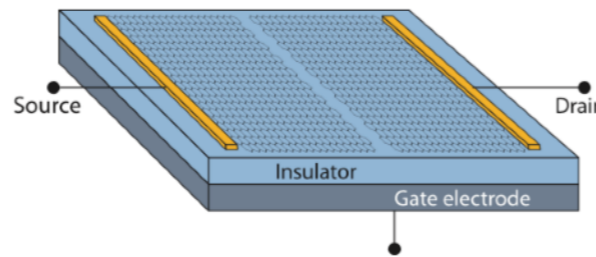


Theoretical studies

- The largest theory lab
- Theory support of experiments : NICA, SHE, neutrino physics, etc.
- Collaboration with: CER, DESY, KEK, Fermilab and many others
- Fellows from 20 countries, 1/3 are young fellows
- Research in:
 - Condense matter
 - Fundamental interactions
 - Elementary particles
 - Nuclear physics
 - Modern math physics

BLTP Publications (2012–2016)

Year	2012	2013	2014	2015	2016	Total
Journal Publications	382	363	364	356	370	1835
Conference Proceedings	97	122	131	142	165	657
Total	479	485	495	498	535	2492



LRB

Laboratory of Radiation Biology

http://lrb.jinr.ru/new/olab/olab_en.shtml

DLNP

Dzheleпов Laboratory of Nuclear Problems

<http://dlnp.jinr.ru>

Structure of DLNP

- Particle Physics
 - Accelerator Technologies
 - Neutrino Physics & Astrophysics
 - Radiation Medicine, Genetics, Molecular Genetics
 - Radiochemistry & Nuclear Spectroscopy
 - IT, design office, workshop, services, etc
 - Education & Outreach
-
- about 650 employees
 - among them about 500 scientific staff

SCIENCE & TECHNOLOGIES

Particle Physics

- ATLAS
- Mu2e, g-2
- COMET
- BES-III
- PANDA

Neutrino Physics & Astrophysics

- BAIKAL GVD
- Daya Bay/JUNO
- NOVA
- BOREXINO
- DANSS
- GERDA
- GEMMA/vGEN
- SuperNEMO
- TUS/Nucleon/TAIGA
- EDELWEISS

Technologies

- Precise Laser Metrology
- New semiconductor detectors
- Ultra cold temperatures

Underlined astrophysical experiments

SCIENCE & TECHNOLOGIES

Medicine & Molecular Genetics

- Proton Therapy
- Medical-biological studies
- Radiation genetics

Education & Outreach

- Schools, conference, seminars
- Web-site of DLNP, social networks
- Lecturing at MSU, MIPT, «Dubna» University and others

Bruno Pontecorvo worked in JINR (1950-1993)
establishing a School of Neutrino Physics



Бруно Понтекорво

Main objective:

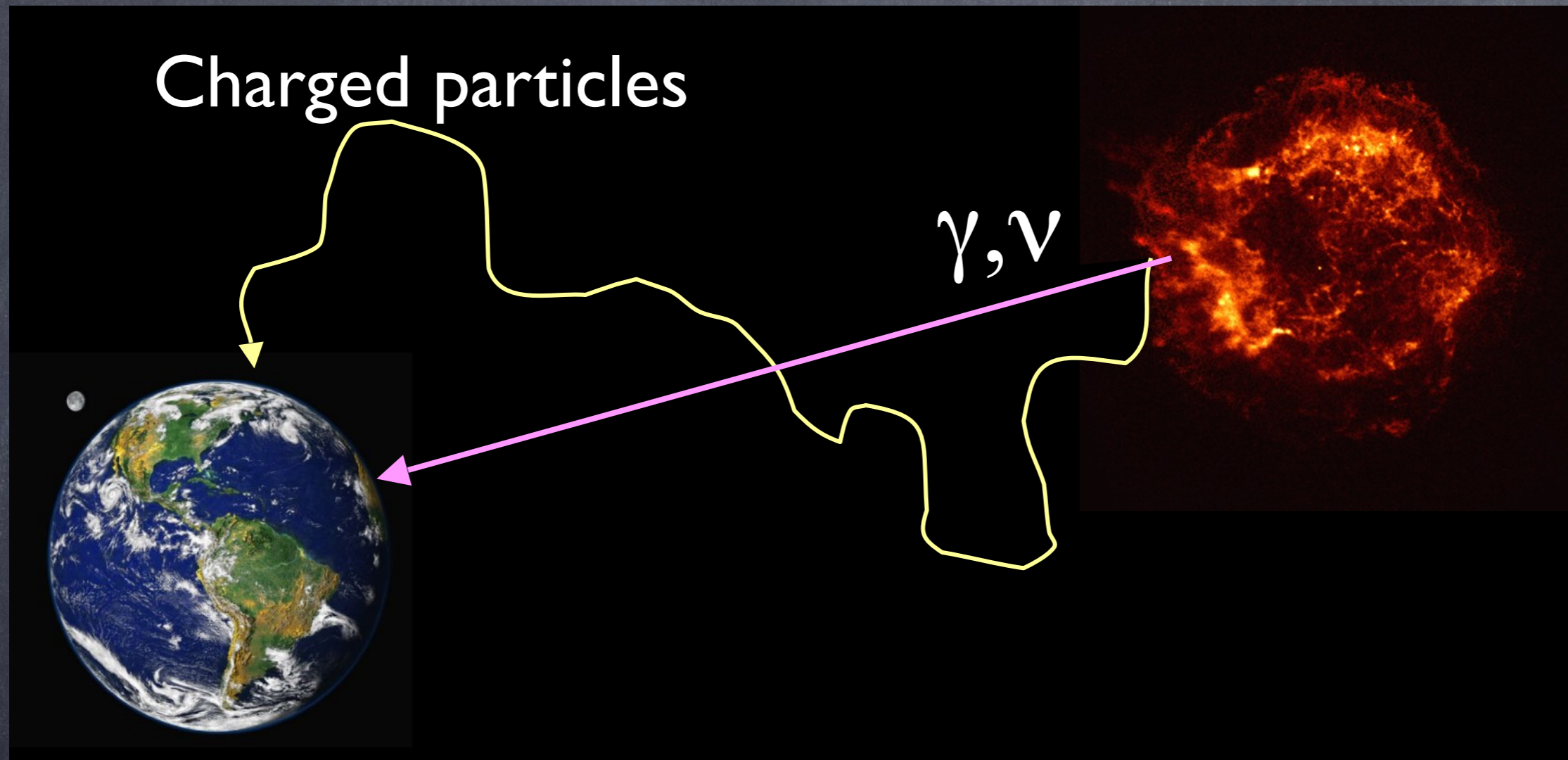
Astrophysical Catastrophes in the Universe

- Sources of ultra-high energy particles



- We are made of star dust
- Modern stars are already of the third generation

Why Neutrino?



- Charged particles loose direction and energy
- Photons get absorbed
- Neutrino astronomy is possible because of weak interaction neutrino

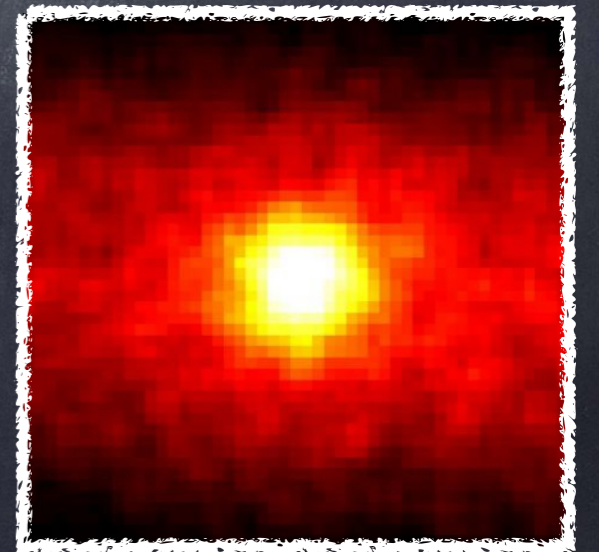
There is no sunrise and sunset with neutrinos



<https://www.youtube.com/watch?v=mu7IYTXP-hI>

Image of Sun with Neutrinos by
SuperKamiokande

Credit: R.Svoboda, K.Gordan (LSU)



Short History of Neutrino Telescopes



M.A. Markov. 1960

«We propose to install detectors deep in a lake or in the sea and determine the direction of charged particles with the help of Cherenkov radiation». ICHEP, Rochester. p578

Short History of Neutrino Telescopes

1960 - M.Markov - main idea.

1976 - Discussions of DUMAND project

1980 - Start of works on construction of BAIKAL Detector lead by G.V.Domogatsky

1993 - NT-36 (36 OM) @BAIKAL

1996 - NT-96 (96 OM)

1997 - AMANDA B10 (302 OM) @South Pole

1998 - NT-200 (192 OM)

2000 - AMANDA II (677 OM)

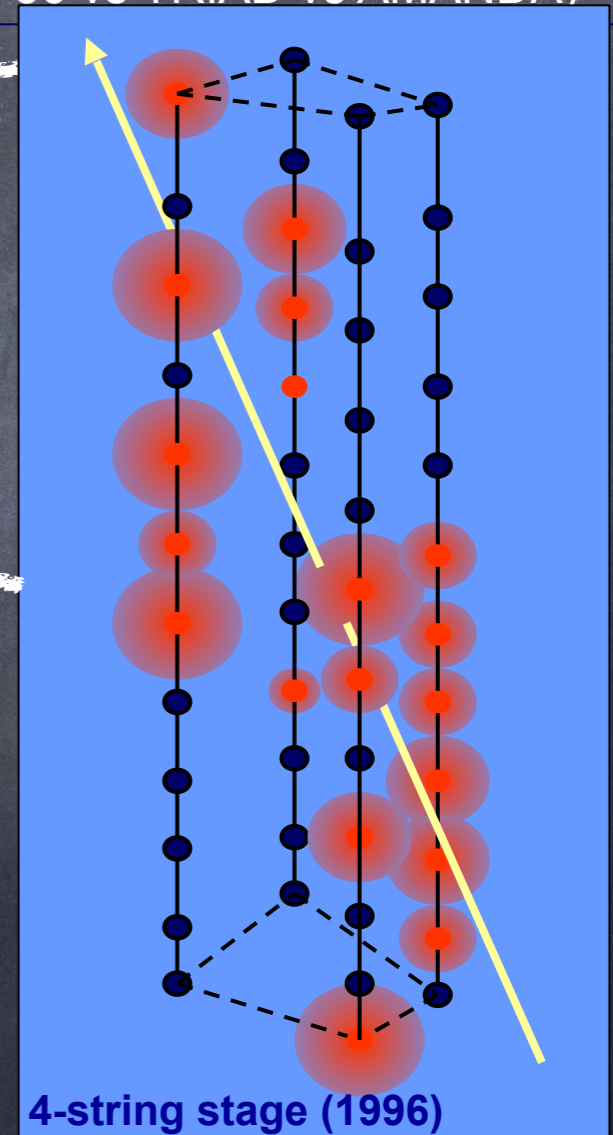
2005 - NT-200+ (228 OM)

2005 - IceCube (first string)

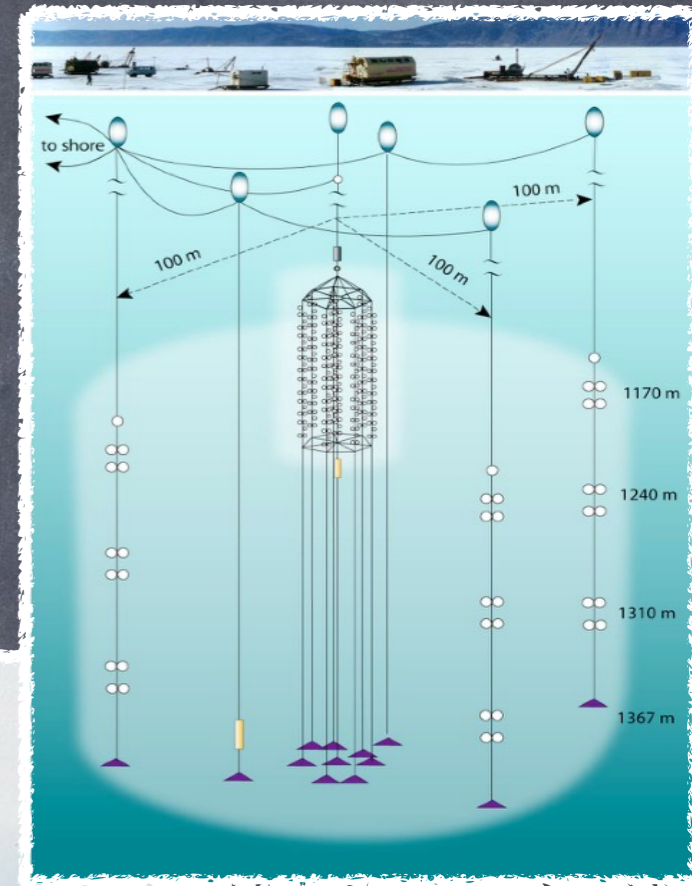
2010 - IceCube (last string)

2015 - BAIKAL GVD («Dubna» cluster)

J. Learned to C.Spiering:
„Congratulations for winning
the 3-string race!“
(NT-36 vs TRIAD vs AMANDA)



BAIKAL Neutrino Telescope Pioneered the field

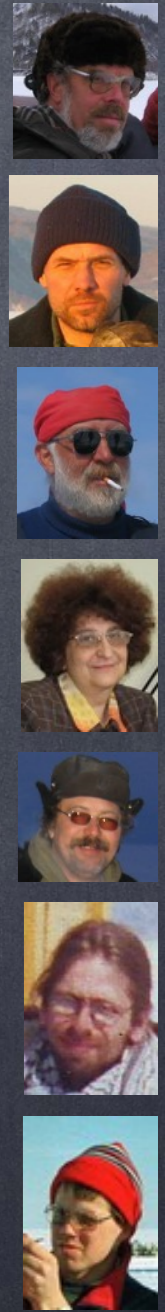


BAIKAL Neutrino Telescope Pioneered the field



DESY was a major collaborator in BAIKAL in 90's

BAIKAL Neutrino Telescope Pioneered the field



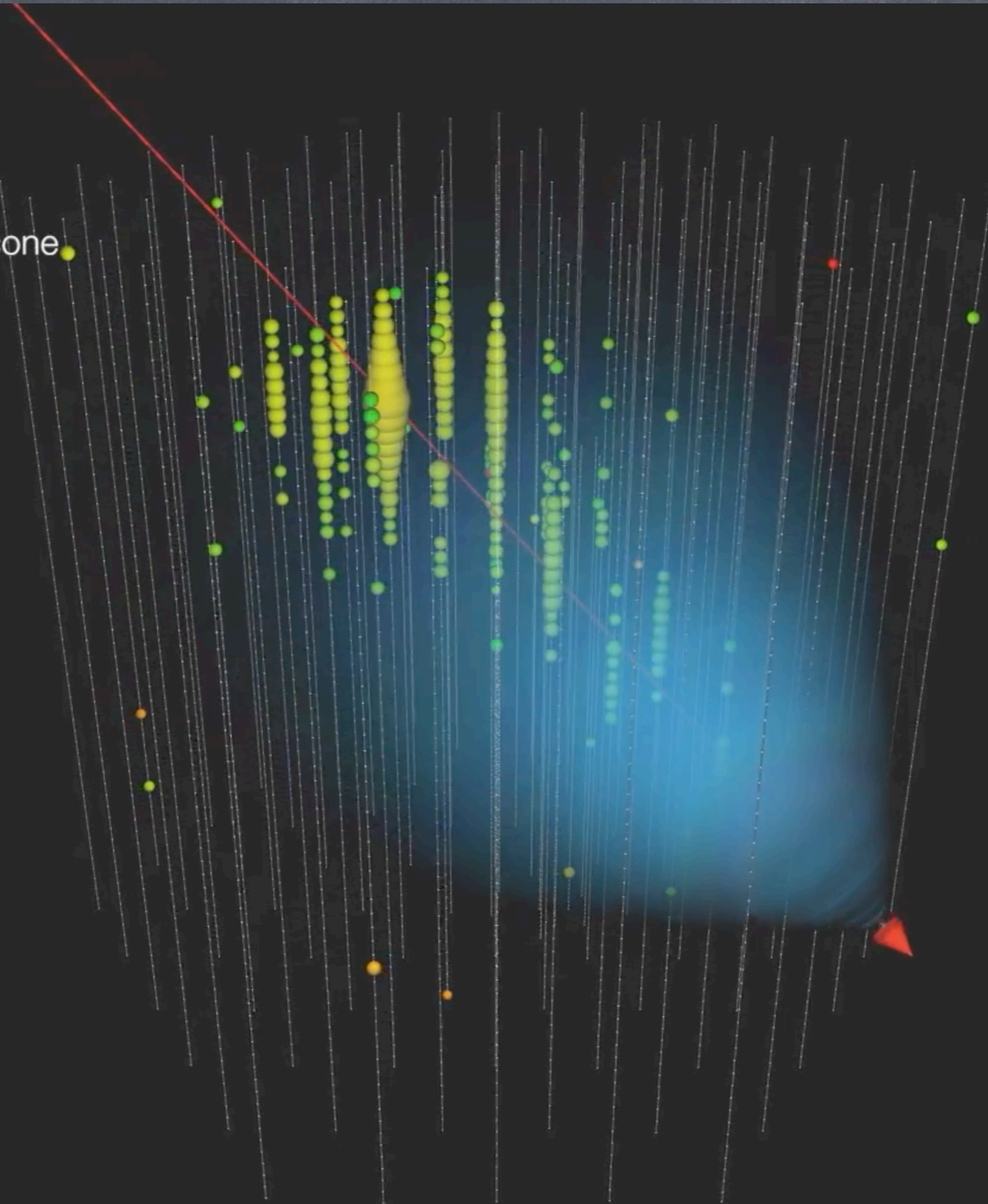
1990

A new boost in BAIKAL Neutrino Telescope history:

- Discovery of UltraHigh energy Neutrinos by IceCube (2014)
- JINR major contribution to construction of cubic-km BAIKAL GVD

IceCube event

with simulated Cherenkov cone



Why BAIKAL?

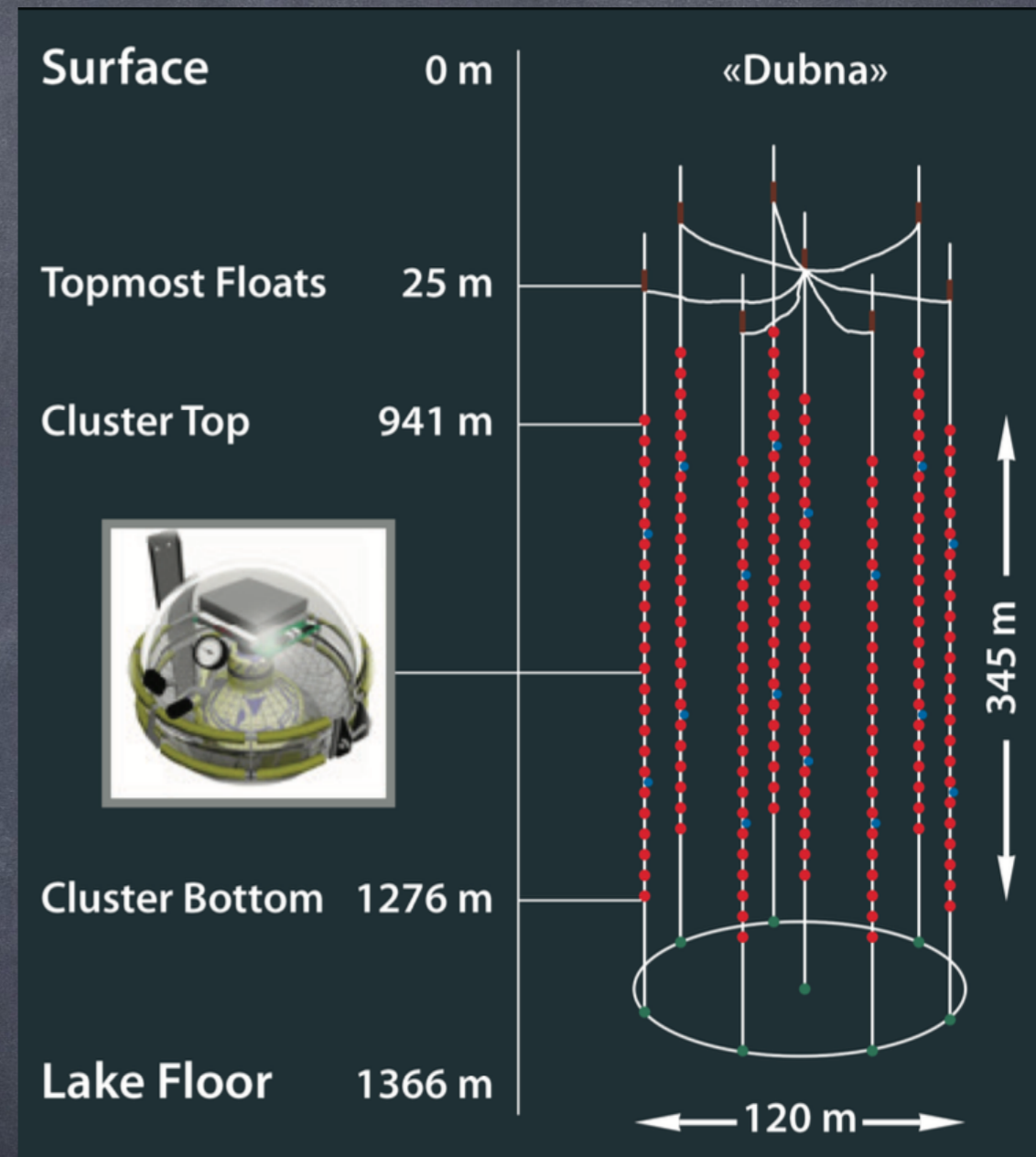
Accurate determination of arrival direction
in BAIKAL water



Light re-scattering in ice is large

The Plan

- Main Goal
 - Point sources of UHE neutrino
 - 3D Array of photo-sensors
 - Phase I: 0.4 km³ (by 2021)
 - Phase II: 1.5 km³ (by 2027)
- Installation site
 - South Baikal
 - Depth 1.4 km
 - Distance from shore 3.5 km
- Requirements
 - Adjustable structure
 - Synchronization < 1ns



Deployment status

	2015	2016	2017	2018
Number of clusters	≤1	1	2	3
Number of OM	192	288	576	864



The diagram illustrates the deployment of the IceCube Neutrino Telescope. It shows three vertical detector strings, each 750 m long, with a diameter of 300 m. The strings are spaced 125 m apart. The detector modules are arranged in a grid pattern. The diagram also shows the Ostankino Tower in Moscow for comparison, which is 540 m tall. The detector strings are shown to be significantly taller than the tower. The diagram also shows the detector strings are 91 m apart at the base. The detector strings are shown to be 525 m apart at the top. The detector strings are shown to be 36 OM apart at the top. The detector strings are shown to be 91 m apart at the base. The detector strings are shown to be 525 m apart at the top. The detector strings are shown to be 36 OM apart at the top.

The largest Neutrino Telescope in Northern hemisphere

Expedition 2019 Ongoing!

The goals:

- Install two more clusters
- Repair part of cluster #3



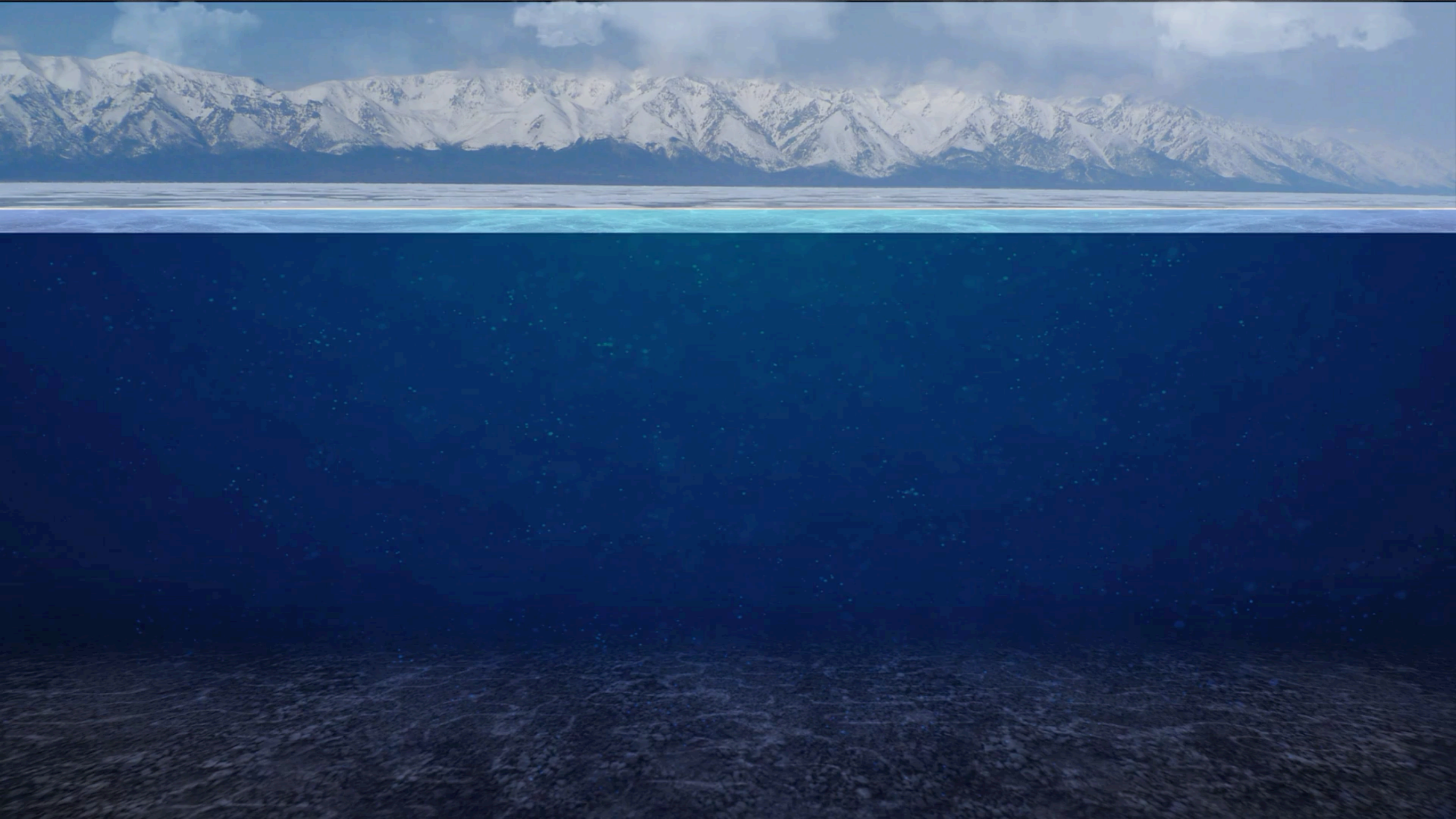
Hello from Baikal.
We are installing
two new clusters



Two new clusters installed in 2019!



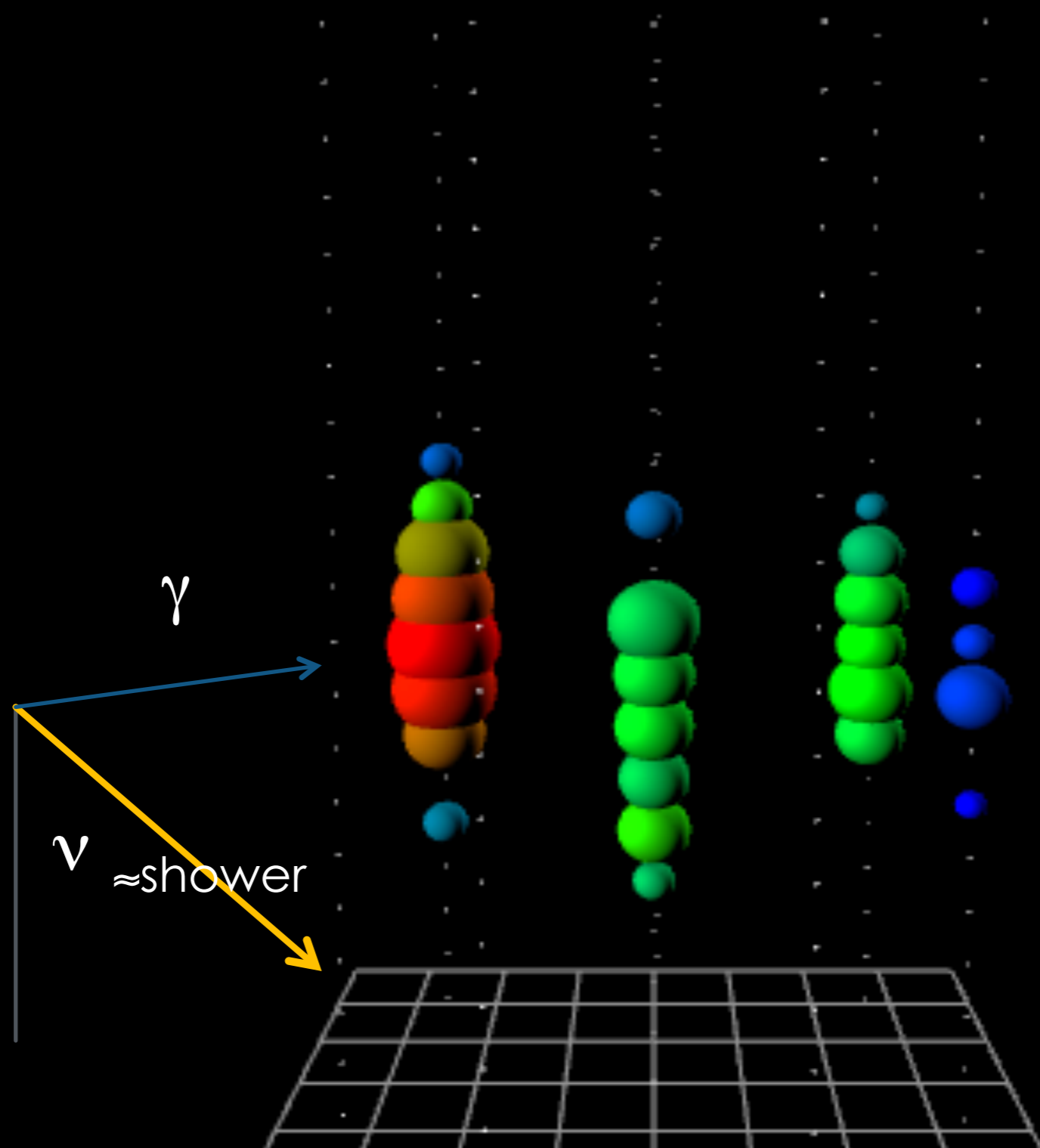
BAIKAL GVD Construction



Physics

Neutrino candidate

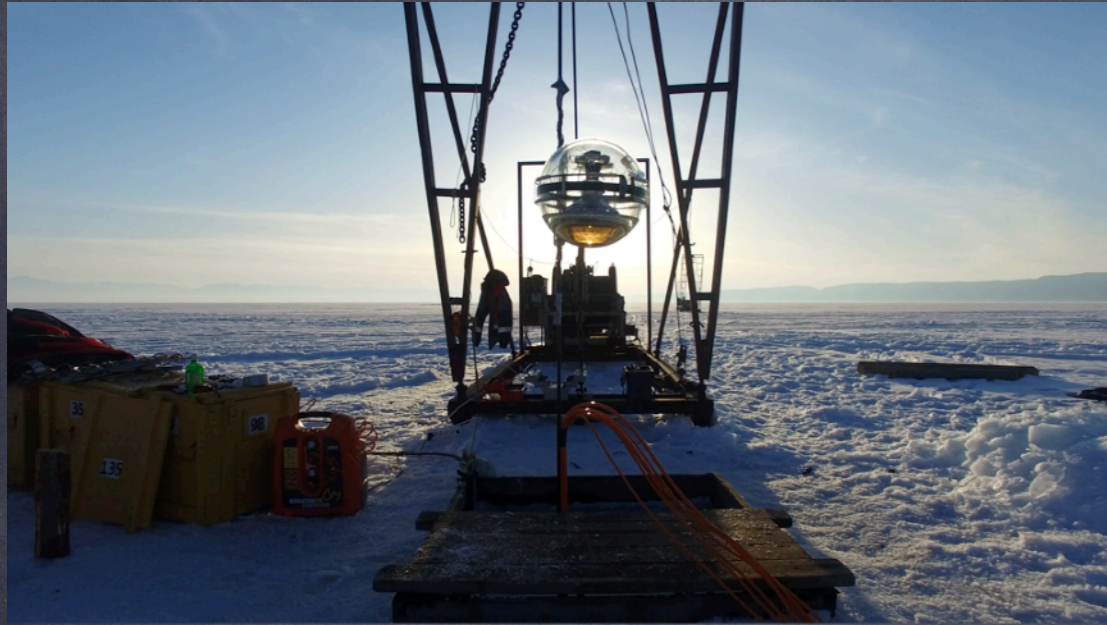
$E = 158 \text{ TeV}$, $\theta = 59^\circ$, $\rho = 73 \text{ m}$, $z = -62 \text{ m}$



Neutrino Astronomy unveiled



The most romantic experiment ever



Women's day!



Way for lunch

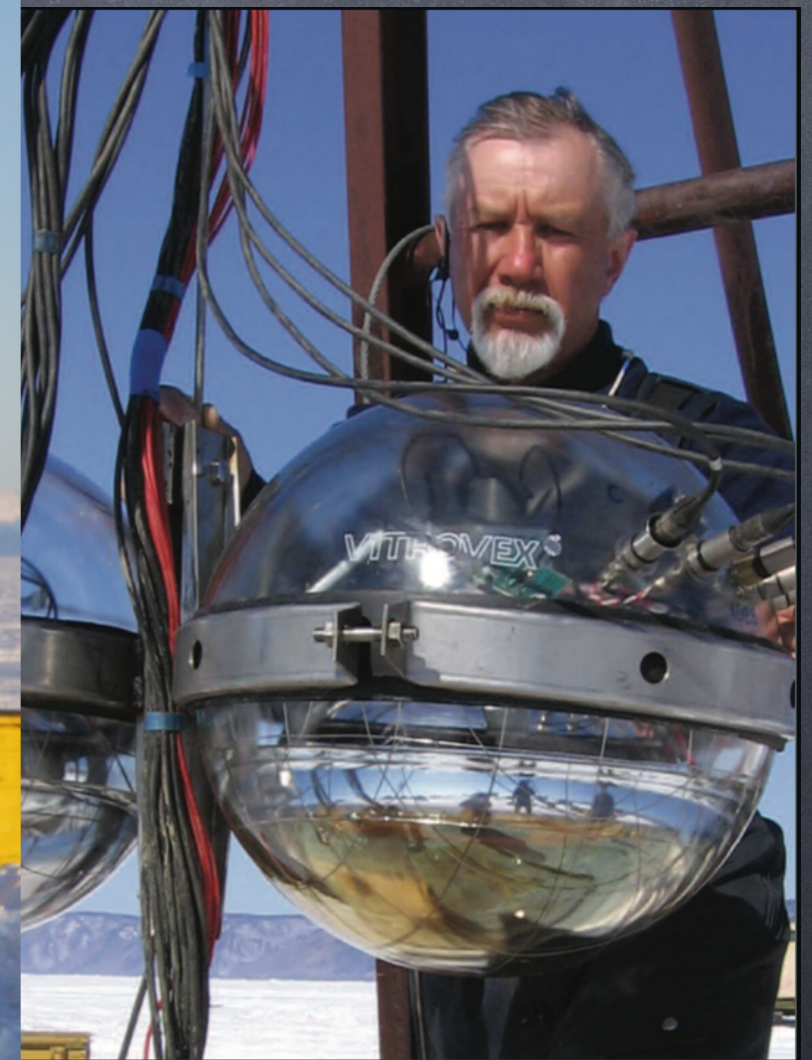


JINR

The End

Additional Materials

Baikal Optical Module (OM)



Canteen

Restore
Houses

Sewage &
Water supply

Plans
@SITE

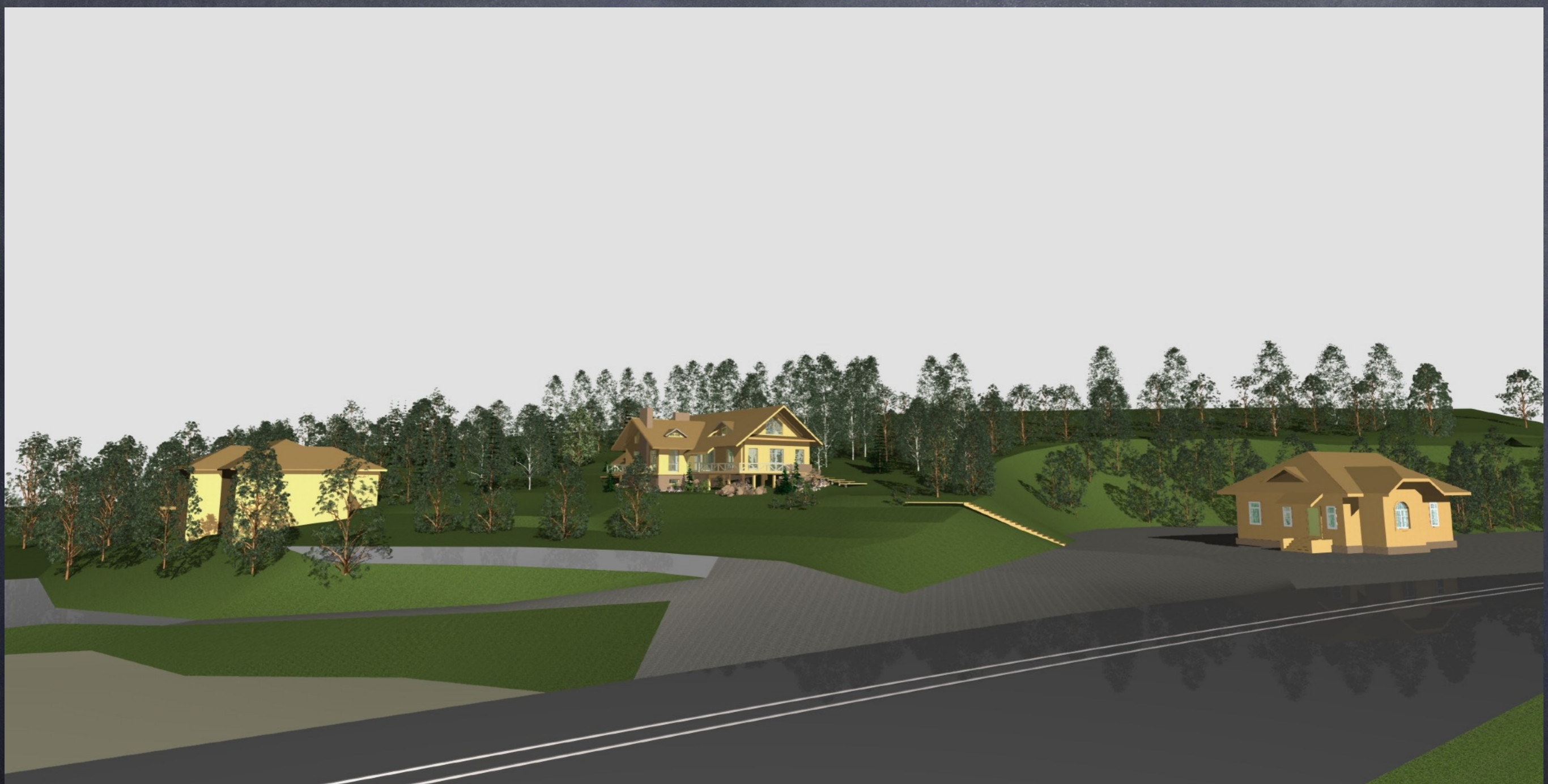
Cable
channel

Renovate
Houses

Pier for
2 boats

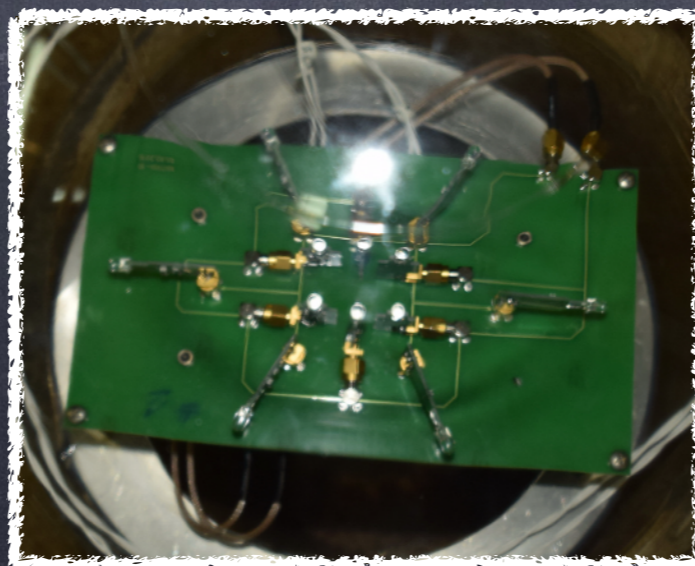
Property

Design view @ 106 km



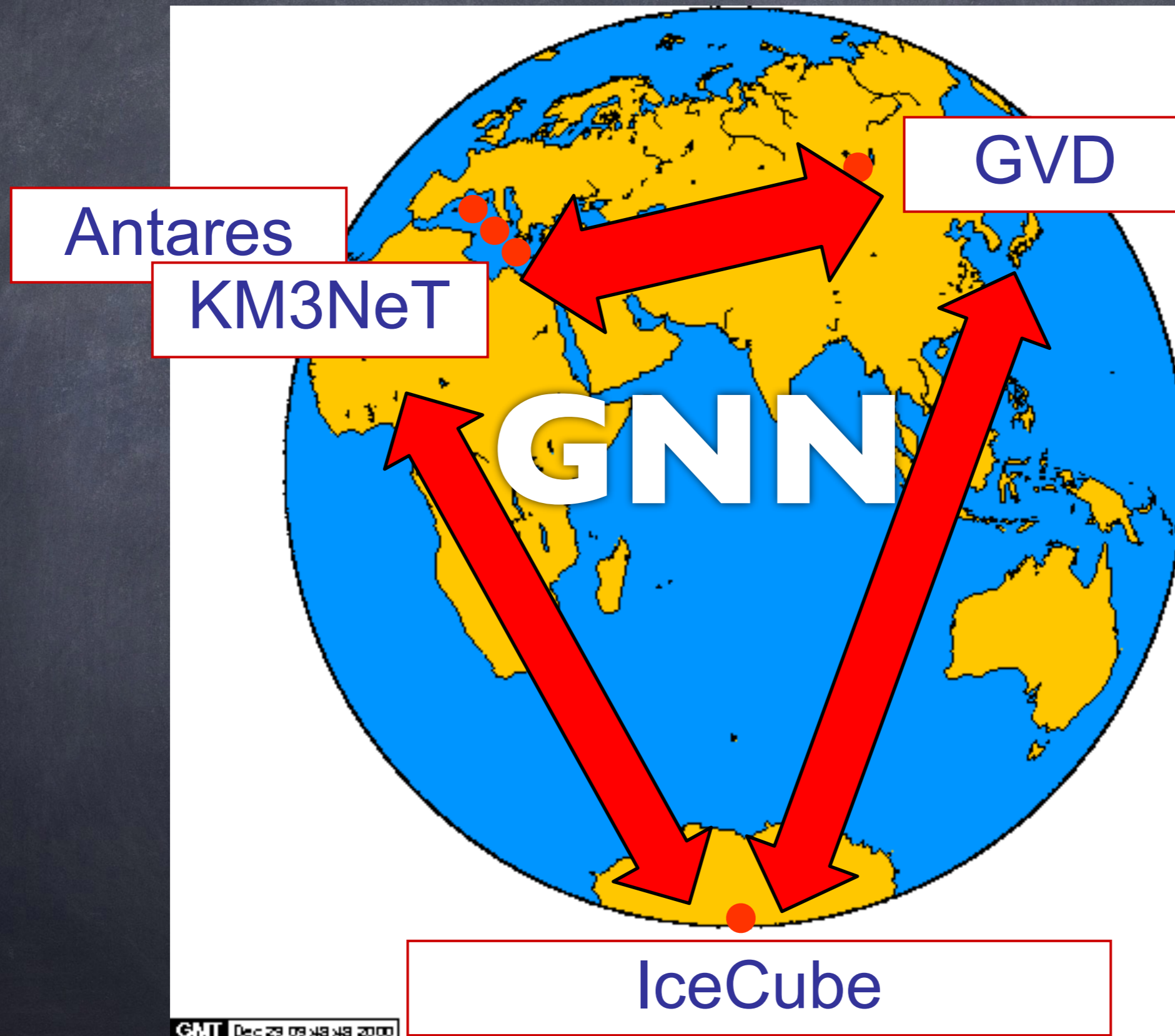
Detector performance

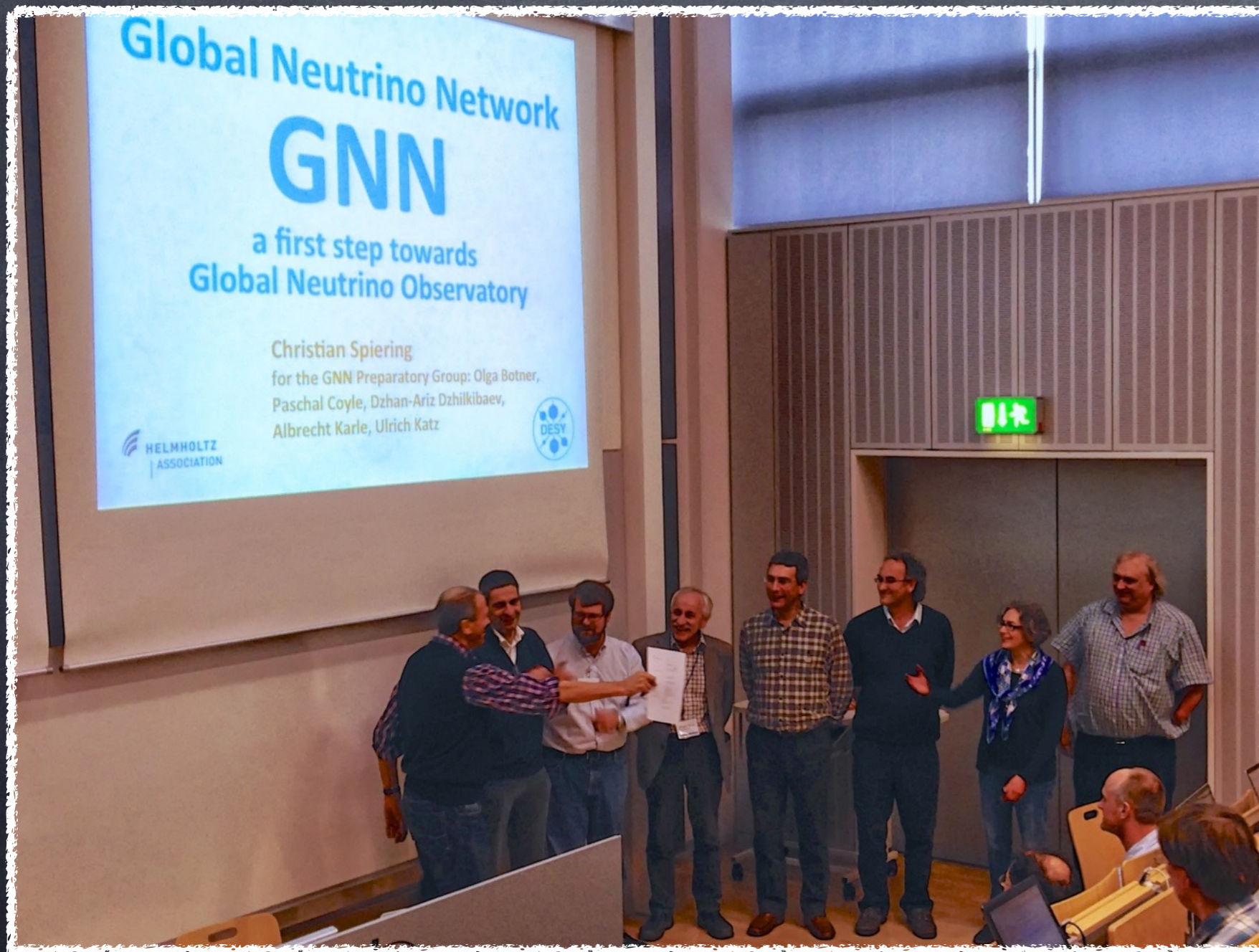
- Position of OMs
 - Acoustic system (few cm)
 - Time synchronization between OMs
 - @same cluster & between clusters
 - Laser
 - LED matrix
- } 1-2 ns



BAIKAL GVD in global context

Global Neutrino Network





Oct. 2013,
Munich

- Antares
- Baikal
- IceCube
- KM3NeT

<http://www.globalneutrino.org/>

Infrastructure upgrade

OM assembling hall @DLNP JINR



12 OM/day

Vacuuming

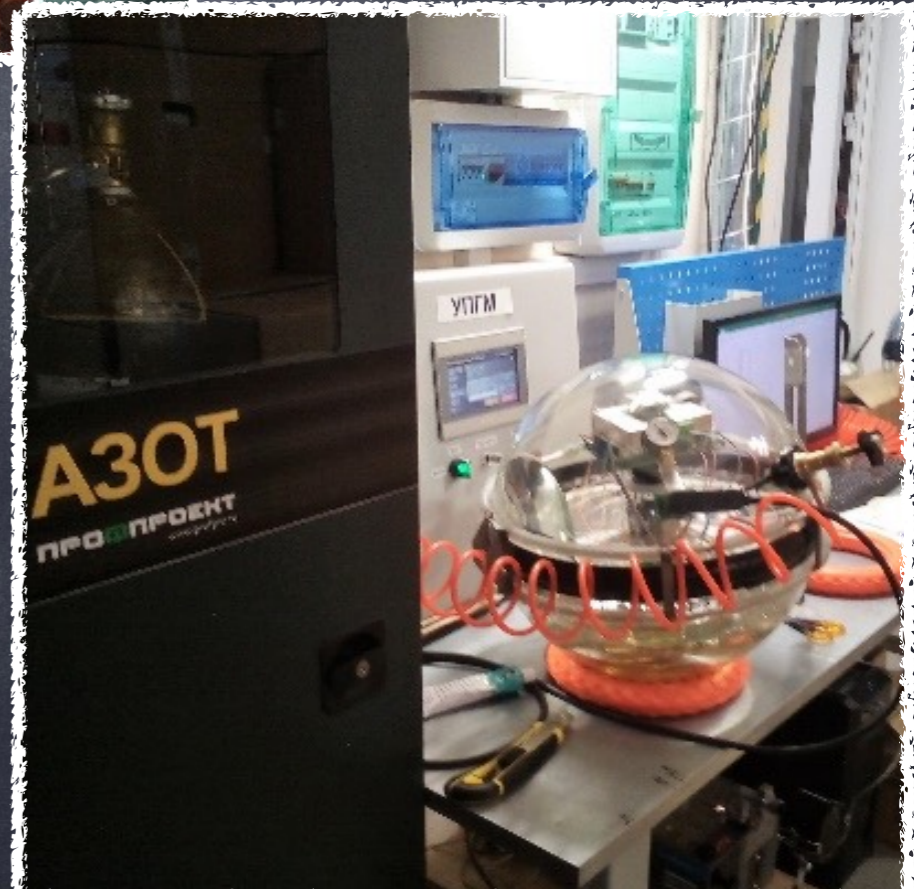


Full testing



JINR facilities

Nitrogen drying



DAQ testing @INR



- Long term tests
- All cluster components
- Full power load

Long-term testing @JINR



@SITE

New data taking center



Purchased storage building @ Baikalsk



New shore lab



New living boxes



@SITE

Transportation issues solved

